

## Disclaimer

The solution at hand was written in the course of the respective class at the University of Bonn. If not stated differently on top of the first page or the following website, the solution was prepared and handed in solely by me, Marvin Zanke. Anything in a different color than the ball pen blue is usually a correction that I or a tutor made. For more information and all my material, check:

<https://www.physics-and-stuff.com/>

**I raise no claim to correctness and completeness of the given solutions! This equally applies to the corrections mentioned above.**

This work by [Marvin Zanke](#) is licensed under a [Creative Commons Attribution-NonCommercial-ShareAlike 4.0 International License](#).

# General Relativity 7. Home exercise

Martin Zank

7/7  
Good

2

05.06.2018

110)  $G(r) = \int \epsilon_0 = \text{const}, r < r_0$   
 $0, r > r_0$

Radius  $r_0$   
 free input  
 parameter -  
 what else?

$g_{00} = e^{\lambda(r)}, g_{11} = -e^{-\lambda(r)}, g_{22} = -r^2, g_{33} = -r^2 \sin^2 \theta$

$G^0_0 = \frac{1}{r^2} - \left( \frac{1}{r^2} - \frac{\lambda'(r)}{r} \right) e^{-\lambda(r)}$

$G^1_1 = \frac{1}{r^2} - \left( \frac{1}{r^2} + \frac{\lambda'(r)}{r} \right) e^{-\lambda(r)}$

$G^2_2 = G^3_3 = -\frac{1}{2} (\nu''(r)) + \frac{\nu'(r)}{2} - \frac{\lambda'(r)\nu'(r)}{2} + \frac{\nu' - \lambda'}{r} e^{-\lambda(r)}$

$T^0_0 = \epsilon, T^1_1 = -p, T^2_2 = -p, T^3_3 = -p, G^i_j = 8\pi G T^i_j$

meaning?

For each  $\epsilon$  and  $p$ , we could choose  $\lambda$  arbitrary s.t.  $G^i_j$  fulfilled

a)

$G^0_0 = 8\pi G T^0_0 \Rightarrow 8\pi G \epsilon_0 = \frac{1}{r^2} - \left( \frac{1}{r^2} - \frac{\lambda'}{r} \right) e^{-\lambda}$

$\Rightarrow 8\pi G \epsilon_0 r^2 = 1 - (1 - \lambda' r) e^{-\lambda} = 1 - e^{-\lambda} + \lambda' r e^{-\lambda}$

$= \frac{d}{dr} (r - r e^{-\lambda})$

$\Rightarrow 8\pi G \int_0^r dp p^2 \epsilon_0 = r - r e^{-\lambda} = r(1 - e^{-\lambda})$

$= \frac{M(r)}{4\pi}$

$\epsilon_0 = (\epsilon + p) \dots$   
 $g_{\mu\nu} \dots = 1$   
 $T^0_0 = \frac{1}{g_{00}} (\epsilon + p)$   
 $T^i_i = (\epsilon + p) - p$   
 also if  $g$  not diagonal?

I would think

so

$\Rightarrow 2GM(r) = r(1 - e^{-\lambda})$

$\Rightarrow e^{-\lambda} = 1 - \frac{2GM(r)}{r}$

$\Rightarrow e^{\lambda} = \left( 1 - \frac{2GM(r)}{r} \right)^{-1}$

$r < r_0: M(r) = 4\pi \int_0^r dp p^2 \epsilon_0 = \frac{4\pi}{3} r^3 \epsilon_0$

$\Rightarrow e^{\lambda} = \left( 1 - \frac{8\pi G r^3 \epsilon_0}{3r} \right)^{-1} = \left( 1 - \frac{r^2}{R^2} \right)^{-1}$

$R^2 = \frac{3}{8\pi G \epsilon_0}$

Why not evaluate  $M_0$  as well?

can be easily cal.

$r > r_0: M(r) = 4\pi \int_0^{r_0} dp p^2 \epsilon_0 \equiv M_0 = \frac{4\pi r_0^3 \epsilon_0}{3}$

$\Rightarrow e^{\lambda} = \left( 1 - \frac{2GM_0}{r} \right)^{-1}$

$\rightarrow \frac{1}{r} \left( \frac{r_0^3}{r^2} \right)$

b) Also had a third eq. arising from  $T_{rr}$  or the Einstein eq.

1

$$-2p'(r) = v'(r)(p(r) + \epsilon(r))$$

$$\Rightarrow p'(r) = -\frac{v'(r)}{2}(p(r) + \epsilon_0) = -\frac{v'(r)}{2}p(r) - \frac{v'(r)}{2}\epsilon_0 \quad (*)$$

$\Rightarrow$  first solve hom. diff. eq.  $p'(r) = -\frac{v'(r)}{2}p(r)$  with:  
 $v' = \frac{-2p'}{\epsilon_0 + p}$   
 $= -2 \log(\epsilon_0 + p)'$

$\Rightarrow p(r) = C e^{-\int v'/2 dr}$

with use the special solution  $p(r) = -\epsilon_0$  for the inhom. diff. eq. and add it to  $p(r)$ .

$\Rightarrow p(r) = C e^{-\int v'/2 dr} - \epsilon_0$  solves  $(*)$

$\Rightarrow p(r) + \epsilon_0 = C e^{-\int v'/2 dr}$

c) From  $G_{00} = 8\pi G T_{00}$  and  $G_{rr} = 8\pi G T_{rr}$ , we find

2

$$8\pi G \epsilon_0 = \frac{1}{r^2} - \left(\frac{1}{r^2} - \frac{\lambda'}{r}\right) e^{-\lambda}$$

$$-8\pi G p = \frac{1}{r^2} - \left(\frac{1}{r^2} + \frac{v'}{r}\right) e^{-\lambda}$$

$\Rightarrow 8\pi G(p + \epsilon_0) = \frac{\lambda' + v'}{r} e^{-\lambda} = 8\pi G c e^{-\lambda/2}$

$\Rightarrow (\lambda' + v') e^{\lambda/2} = 8\pi G c r e^{\lambda}$

$\Rightarrow v' e^{\lambda/2} = 8\pi G c r e^{\lambda} - \lambda' e^{\lambda/2}$

$y = e^{\lambda/2} \Rightarrow 2y' = 8\pi G c r e^{\lambda} - \lambda' y$

Use the ansatz  $y = e^{\lambda/2} = A - B(1 - \frac{r^2}{r^2})^{1/2} \Rightarrow y' = \frac{B\lambda'}{2} e^{-\lambda/2} = A - B e^{-\lambda/2}$

$\Rightarrow B\lambda' e^{-\lambda/2} = 8\pi G c r e^{\lambda} - A\lambda' + B\lambda' e^{-\lambda/2}$

$\Rightarrow 8\pi G c r e^{\lambda} = A\lambda' \Rightarrow c = \frac{A\lambda'}{8\pi G r} e^{-\lambda}$

$\left| e^{-\lambda} = \left(1 - \frac{r^2}{r^2}\right) \Rightarrow -\lambda' e^{-\lambda} = -\frac{2r}{r^2} \Rightarrow \lambda' e^{-\lambda} = \frac{2r}{r^2}\right.$

$\Rightarrow c = \frac{2A}{8\pi G r^2}$

d) From b)  $p(r) = ce^{-\sqrt{r}} - \epsilon_0$

From a)  $p^2 = \frac{3}{8\pi G \epsilon_0} \Leftrightarrow \epsilon_0 = \frac{3}{8\pi G R^2}$

(1)

Using c), we find

$$p(r) = ce^{-\sqrt{r}} - \epsilon_0 = \frac{c}{A - B\sqrt{1 - \frac{r^2}{R^2}}} - \frac{3}{8\pi G R^2}$$

$$= \frac{8\pi G R^2 c - 3A + 3B\sqrt{1 - \frac{r^2}{R^2}}}{8\pi G R^2 \left\{ A - B\sqrt{1 - \frac{r^2}{R^2}} \right\}}$$

c from d)

$$= \frac{3B\sqrt{1 - \frac{r^2}{R^2}} - A}{A - B\sqrt{1 - \frac{r^2}{R^2}}}$$

e) To fix A and B, we use  $p(r) = 0$  and  $e^{\sqrt{r}} = 1 - \frac{2GM_0}{r_0}$

(1)

i.e. continuity of  $p$  and  $e^{\sqrt{r}}$  at the boundary of the star, as from the lecture

$$e^{\sqrt{r}} = \left(1 - \frac{2GM_0(r)}{r}\right) \text{ for } r > r_0$$

$$p(r) = 0 \text{ for } r = r_0$$

$$\Rightarrow p(r_0) = \frac{1}{8\pi G R^2} \frac{3B\sqrt{1 - \frac{r_0^2}{R^2}} - A}{A - B\sqrt{1 - \frac{r_0^2}{R^2}}} \stackrel{!}{=} 0$$

$$\Leftrightarrow A = 3B\sqrt{1 - \frac{r_0^2}{R^2}} = 3B e^{-\lambda(r_0)/2}$$

$$e^{\sqrt{r_0}/2} = A - B\sqrt{1 - \frac{r_0^2}{R^2}} \stackrel{!}{=} \sqrt{1 - \frac{2GM_0}{r_0}}$$

$$= A - \frac{A}{3} = \frac{2A}{3}$$

$$\Leftrightarrow A = \frac{3}{2} \sqrt{1 - \frac{2GM_0}{r_0}}$$

$$\Rightarrow B = \frac{A}{3\sqrt{1 - \frac{r_0^2}{R^2}}} = \frac{\sqrt{1 - \frac{2GM_0}{r_0}}}{2\sqrt{1 - \frac{r_0^2}{R^2}}}$$

$p(r_0) = 0$  as well on the boundary itself?  
 $\Rightarrow$  we don't know exactly what it's at the boundary (microscopically) but macroscopically, take  $p(r_0 + \delta)$  for  $\delta \rightarrow 0$ , which can be assumed to vanish

Can those A, B be simplified further?  
 $\Rightarrow$  not in tutorial